

APEX at the QSO MUSEUM: Molecular gas reservoirs associated with $z \sim 3$ quasars and their link to the extended Ly α emission

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Optical surveys routinely report the detection of cool gas ($T \sim 10^4$ K) reservoirs (as traced mainly by Lyman-alpha (Ly α) emission) extending on halo scales around $z \sim 3$ quasars (e.g., [1], [2], [3] and [4]), but little or nothing is known about the molecular gas reservoir surrounding the targeted objects. In this poster, I will present an APEX/SEPIA180 spectroscopic survey (~ 200 hours, [5]) of the CO(6-5), CO(7-6) and [CI](2-1) emission lines for a sample of nine $z \sim 3$ quasars from the QSO MUSEUM survey [6], which have similar UV luminosities, but very diverse Ly α nebulae. Using these CO and [CI] observations, we constrain the molecular gas masses to be $M_{\text{H}_2} = (0.4-6.9) \times 10^{11} M_{\odot}$ for the detected sources, and $M_{\text{H}_2} < 1.1 \times 10^{11} M_{\odot}$ for non-detections. We discuss their link with the Ly α properties on larger scales. In particular, we find large velocity shifts (from ~ -400 to $\sim +1200$ km s⁻¹) of the Ly α nebulae with respect to the now available molecular redshifts in five sources, suggesting turbulent inflows/outflows around these quasars. We also find that the two most massive molecular reservoirs in our sample are associated with the dimmest and smallest Ly α nebulae. From this, we speculate that obscuration from the host galaxy, due to physical properties or geometry, could reduce the escape of ionizing and Ly α photons emitted by the quasar, ultimately reducing the emission from the cool CGM.

References

- [1] Borisova et al., ApJ, 831, 39 (2016)
- [2] Arrigoni Battaia et al., A&A, 631, 18 (2019)
- [3] Cai et al., ApJS, 245, 23 (2019)
- [4] Farina et al., ApJ, 887, 196 (2019)
- [5] Muñoz-Elgueta et al., accepted for publication in MNRAS (2022)
- [6] Arrigoni Battaia et al., MNRAS, 483, 3162 (2019)