

# The chemical footprint of AGN feedback in the outflowing circumnuclear disk of NGC 1068.

K. -Y. Huang<sup>1</sup>, S. Viti<sup>1,2</sup>, J. Holdship<sup>1</sup>, S. García-Burillo<sup>3</sup>, K. Kohno<sup>4</sup>, A. Taniguchi<sup>5</sup>, S. Martín<sup>6,7</sup>, R. Aladro<sup>8</sup>, A. Fuente<sup>3</sup>, M. Sánchez-García<sup>9</sup>

<sup>1</sup> *Leiden Observatory, Leiden University, PO Box 9513, 2300 RA Leiden, The Netherlands*

<sup>2</sup> *Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK*

<sup>3</sup> *Observatorio Astronómico Nacional (OAN-IGN)-Observatorio de Madrid, Alfonso XII, 3, 28014-Madrid, Spain*

<sup>4</sup> *Institute of Astronomy, The University of Tokyo, Osawa, Mitaka, Tokyo 181-0015, Japan*

<sup>5</sup> *Division of Particle and Astrophysical Science, Graduate School of Science, Nagoya University, Furocho, Chikusa-ku, Nagoya, Aichi 464-8602, Japan*

<sup>6</sup> *European Southern Observatory, Alonso de Córdova, 3107, Vitacura, Santiago 763-0355, Chile*

<sup>7</sup> *Joint ALMA Observatory, Alonso de Córdova, 3107, Vitacura, Santiago 763-0355, Chile*

<sup>8</sup> *Max-Planck-Institut für Radioastronomie, Auf dem Hügel 69, D-53121 Bonn, Germany*

<sup>9</sup> *Centro de Astrobiología (CSIC/INTA), Ctra de Torrejón a Ajalvir, km 4, 28850 Torrejón de Ardoz, Madrid, Spain*

In the nearby (D=14 Mpc) AGN-starburst composite galaxy NGC 1068, it has been found that the molecular gas in the CND is outflowing, which is a manifestation of ongoing AGN feedback (García-Burillo et al. 2014). The induced interaction between the AGN ionized wind & jet with the molecular gas on the CND has produced large-scale molecular shocks on spatial scales of up to 400pc from the AGN. The outflowing gas has a large span of velocities, which likely drive different shock chemistry signatures at different locations in the CND. In this talk we are presenting our recent ALMA multi-line molecular study (Huang et al. 2022) using SiO and HNCO as tracers of chemical differentiation across the CND. With a radiative transfer analysis coupled with Bayesian inference processes, we are able to determine the gas properties of the potentially shocked gas in the CND.

[1] García-Burillo, S., Combes, F., Usero, A., et al. 2014, *A&A*, 567, A125

[2] Huang K.-Y., Viti S., Holdship J., García-Burillo S., Kohno K., Taniguchi A., Martín S., et al., 2022, arXiv, arXiv:2202.05005 (Accepted for publication in *A&A*)